

# Constraining TTL Thermal Structure

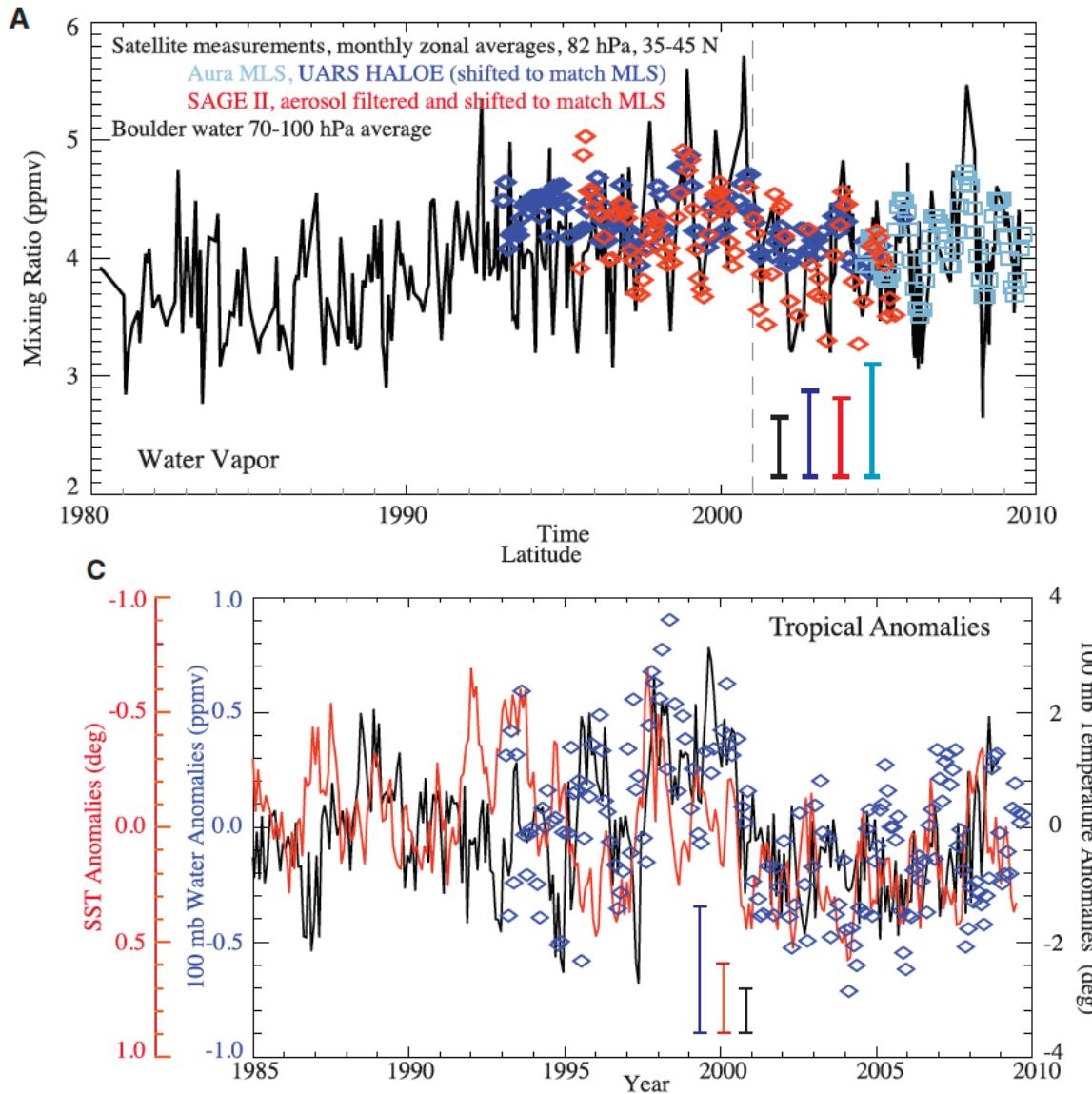
A. Gettelman  
NCAR



# Outline

- Motivation
- Model Simulations
- What determines tropopause temperature
- What is the balance in models?
- What is needed/How Can ATTREX help

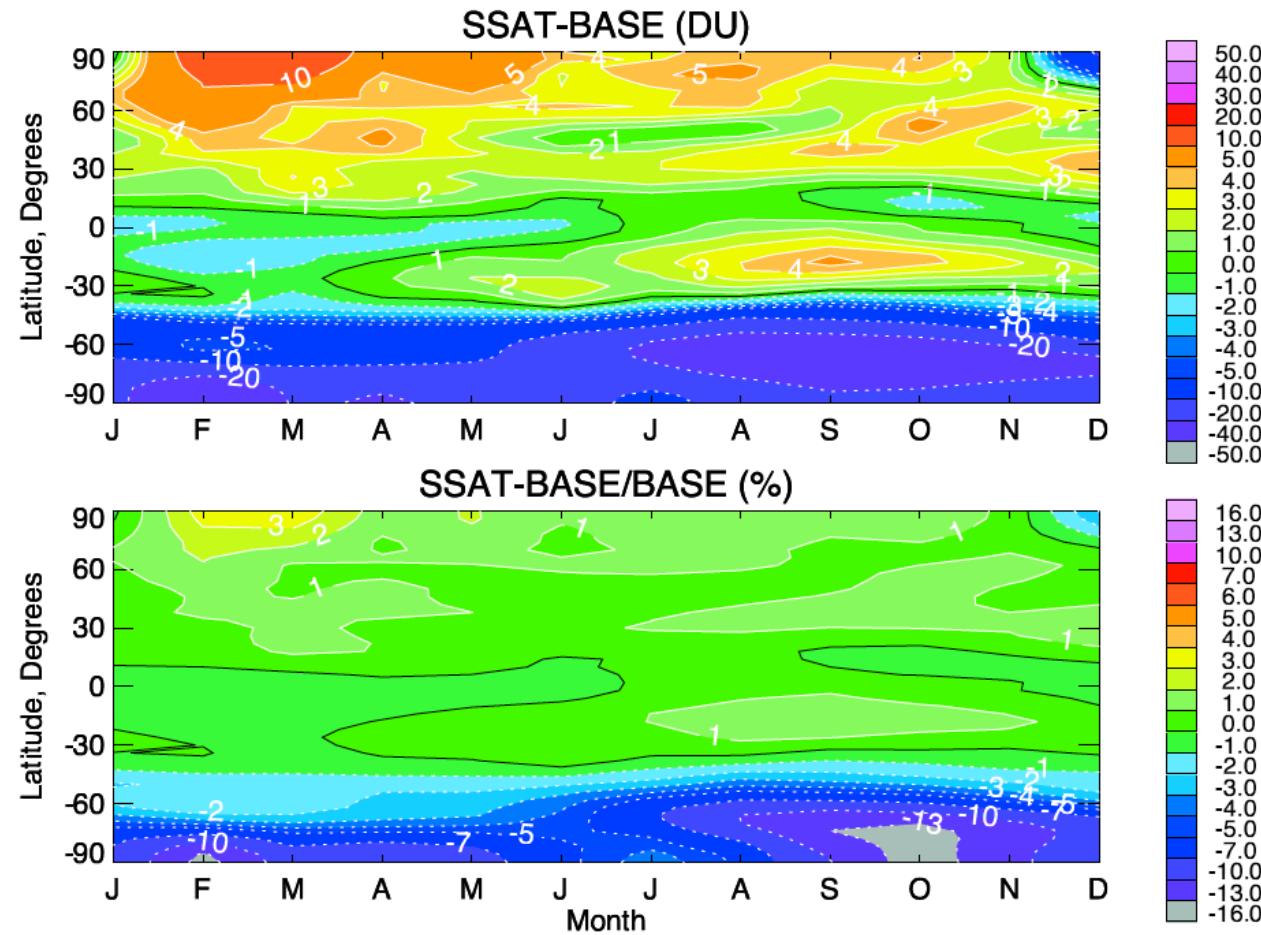
# TTL Temp determines Strat H<sub>2</sub>O



Top: Observed H<sub>2</sub>O  
Bottom: Tropical  
Temp, H<sub>2</sub>O, SST

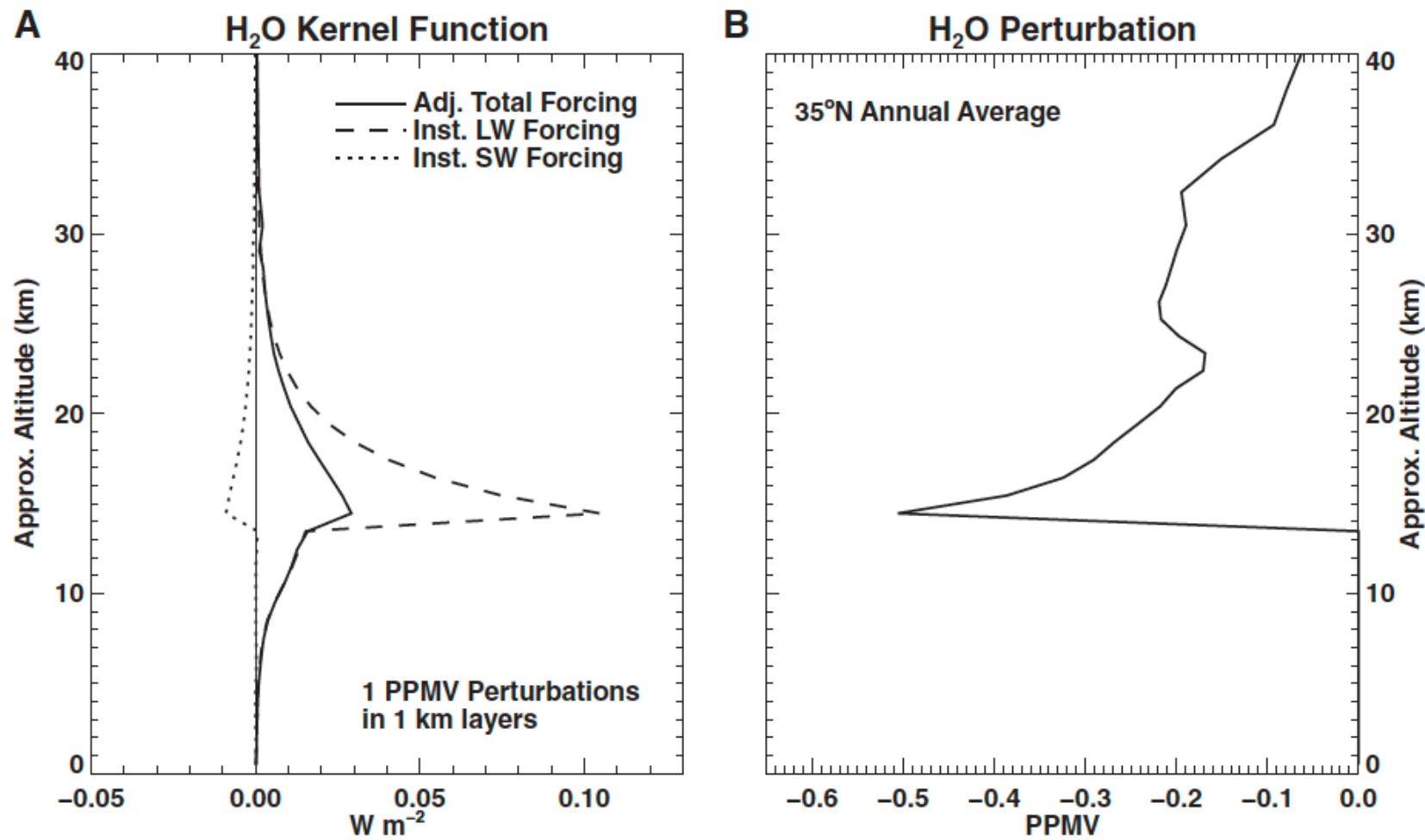
Solomon et al 2010

# Strat H<sub>2</sub>O Affects Chemistry



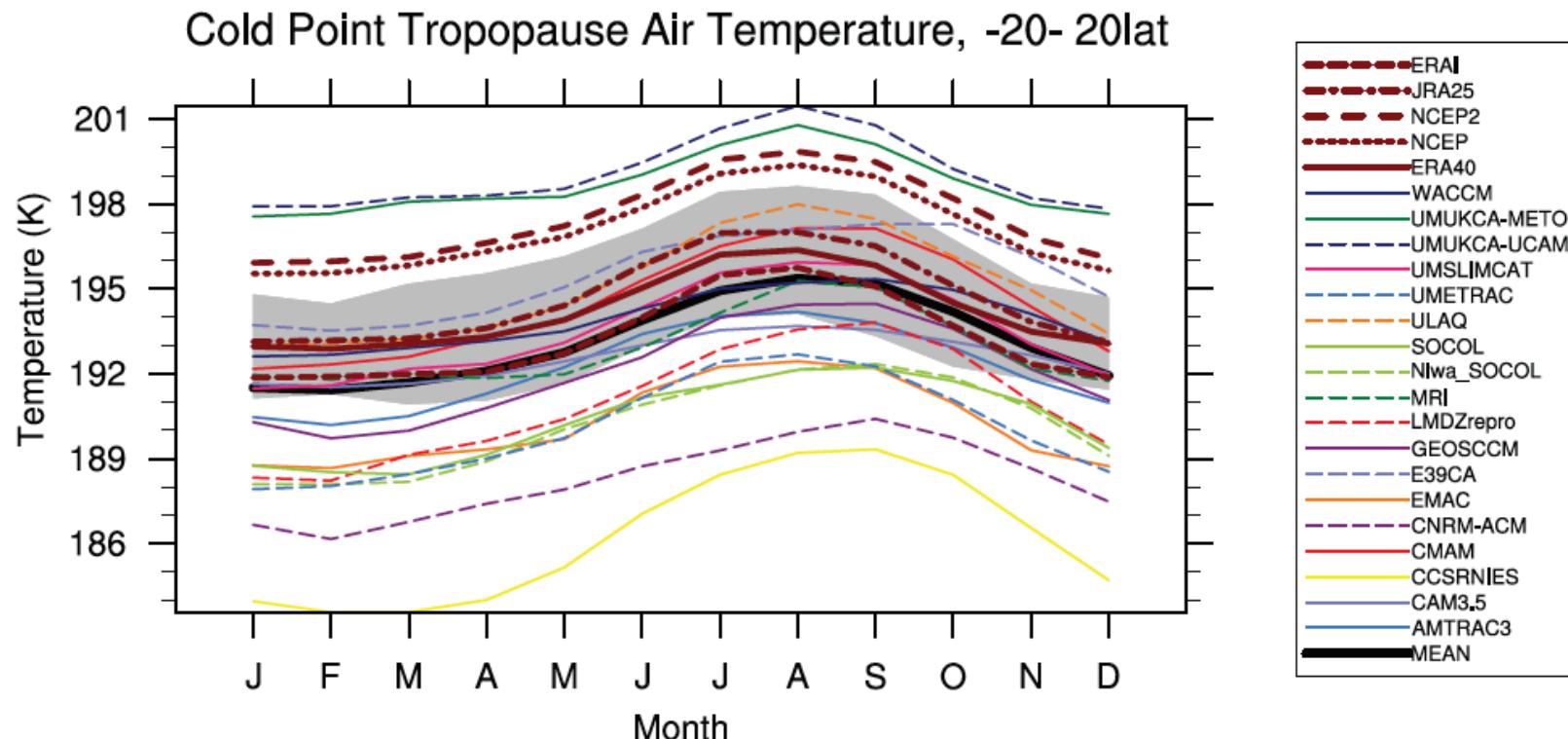
Effect of ~20% increase in H<sub>2</sub>O by allowing supersaturation (Gettelman & Kinnison 2007)  
(combination of a temperature & chemistry effect)

# UTLS H<sub>2</sub>O Affects Radiative Forcing



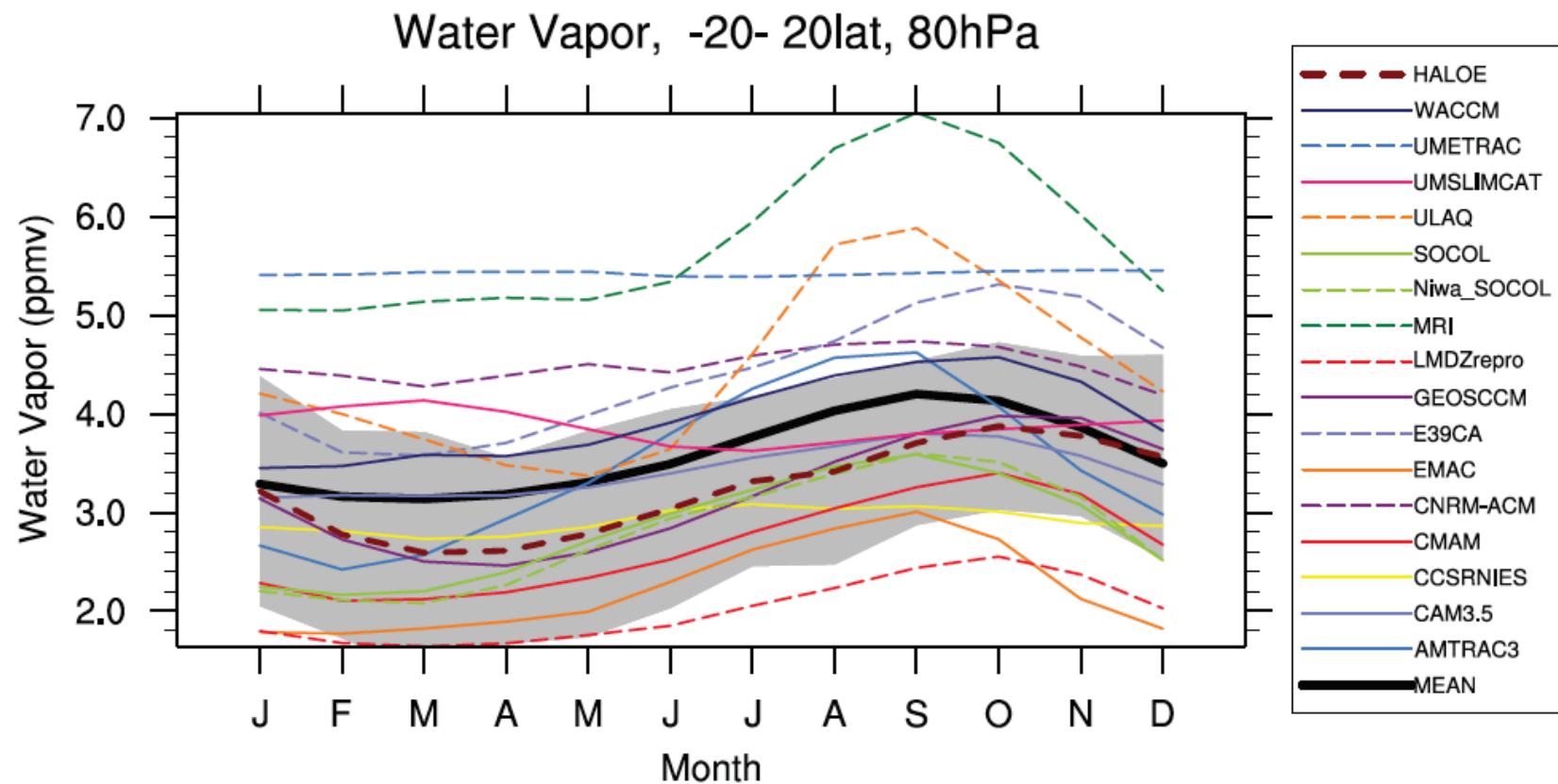
Solomon et al 2010

# Simulating TTL Thermal Structure



SPARC Assessment of CCMs, Ch7 & Gettelman et al, 2010 (JGR)  
Throw out outliers and spread is still ~6K in the mean!  
Note: this is better than CCMVal 1: ~10K.

# Simulated LS Water Vapor

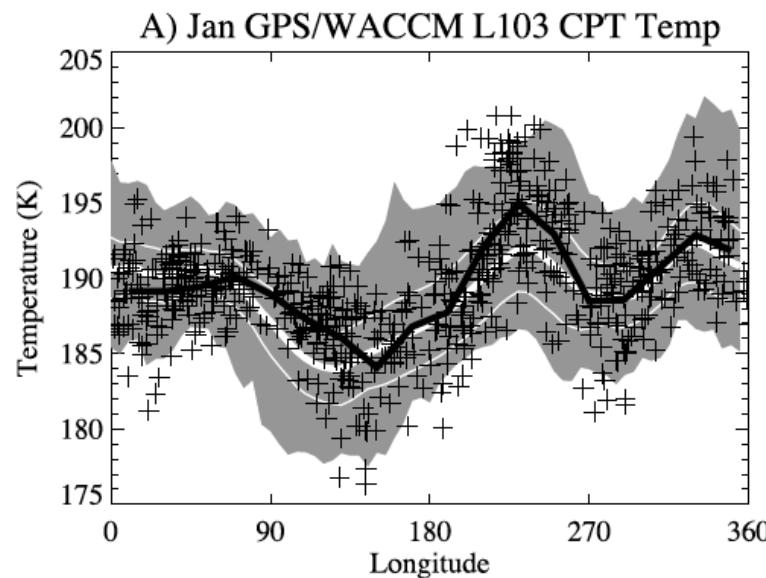


SPARC Assessment of CCMs, Ch7 & Gettelman et al, 2010 (JGR)  
Without outliers, generally within HALOE range.  
Note seasonal shift by ~2 months

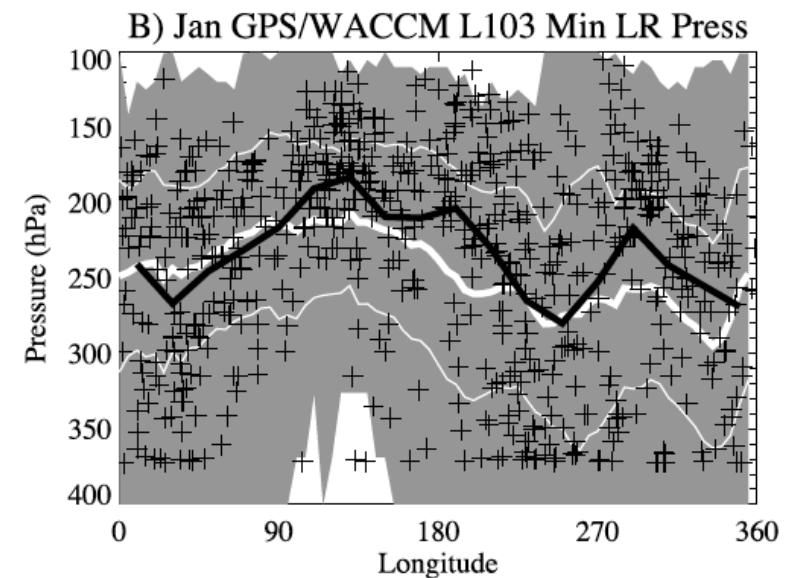
# TTL Temp Variability

Reproduce Large scale (>200km) High Frequency Variability

TTL Top

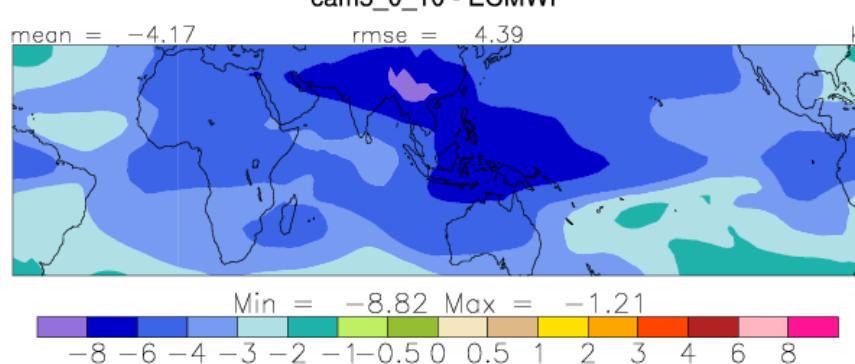
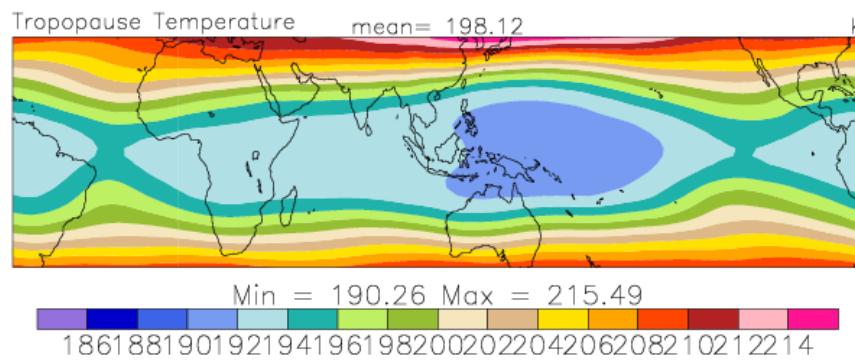
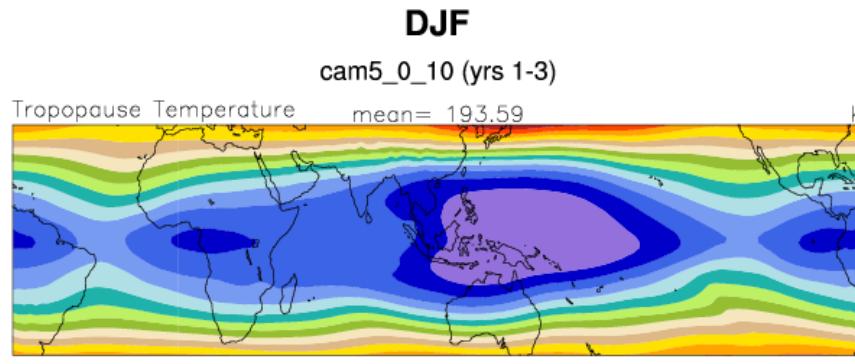


TTL Bottom



Gray: Model  $2\sigma$ , white model mean. Black Crosses and Line: GPS RO temperatures  
Gettelman & Birner, 2007

# CAM5: Latest Climate Model



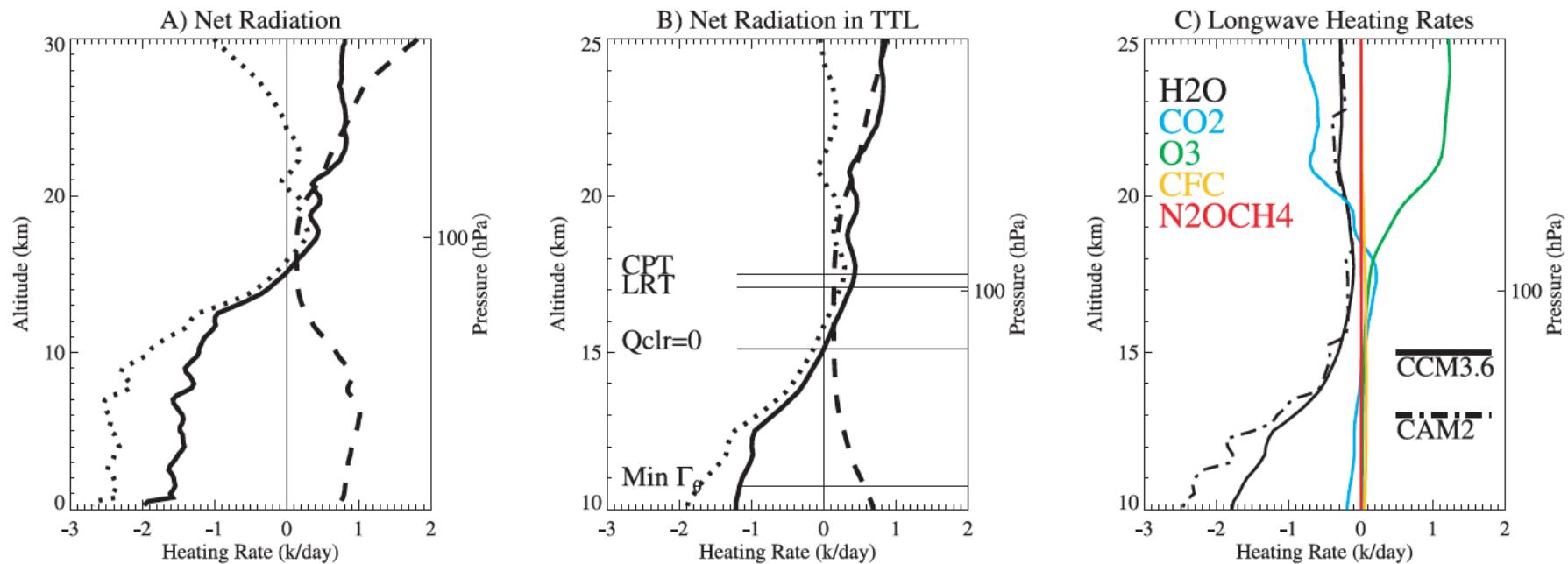
- Seasonal mean ~4K colder than ECMWF reanalyses (DJF)
- Ann mean is ~3.5K cold
- Model is ‘un-optimized’ for CPT
- Need to optimize: hopefully for the right reasons!

# What determines Tropopause Temp?

- Radiative convective equilibrium
- Dynamics ‘creates’ the TTL
  - Brewer-Dobson circulation
- Radiation helps ( $\text{CO}_2$ ): Thuburn & Craig 2002
- Cloud heating also helps
  - Biggest uncertainty...

# Radiative Equilibrium

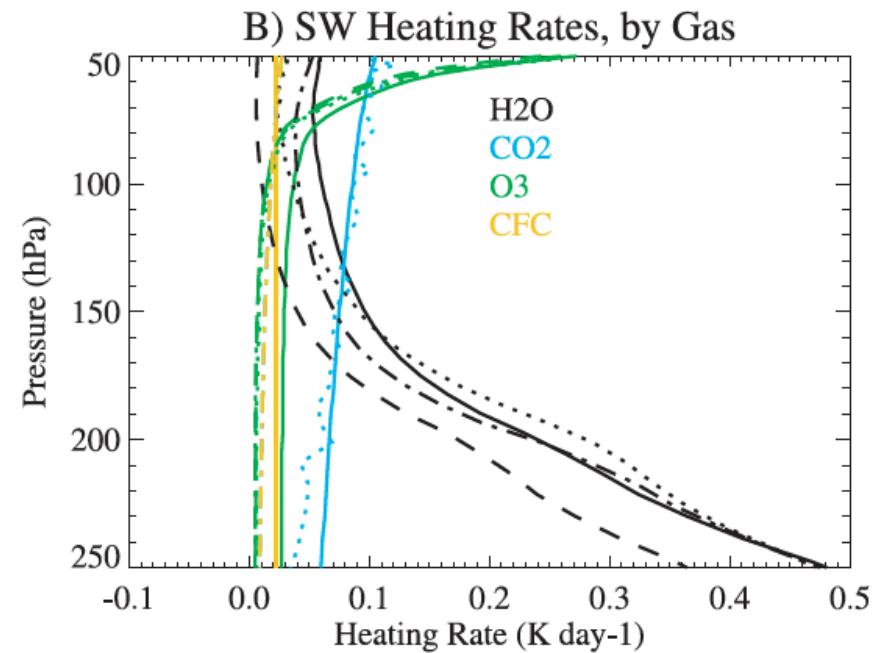
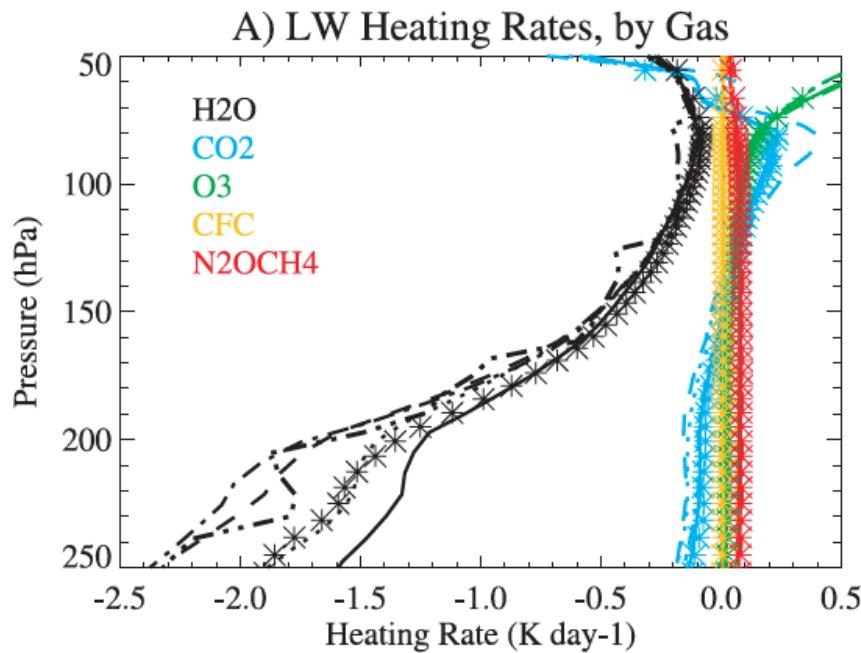
- Heating Rates: Gettelman et al 2004



TTL is a balance of LW cooling ( $H_2O$ ) then heating from  $O_3$  (SW & LW)  
At CPT:  $CO_2$  LW heating does matter

# Radiative Equilibrium (2)

- Heating Rates: Gettelman et al 2004  
Colors=Gases Linestyles=Different models

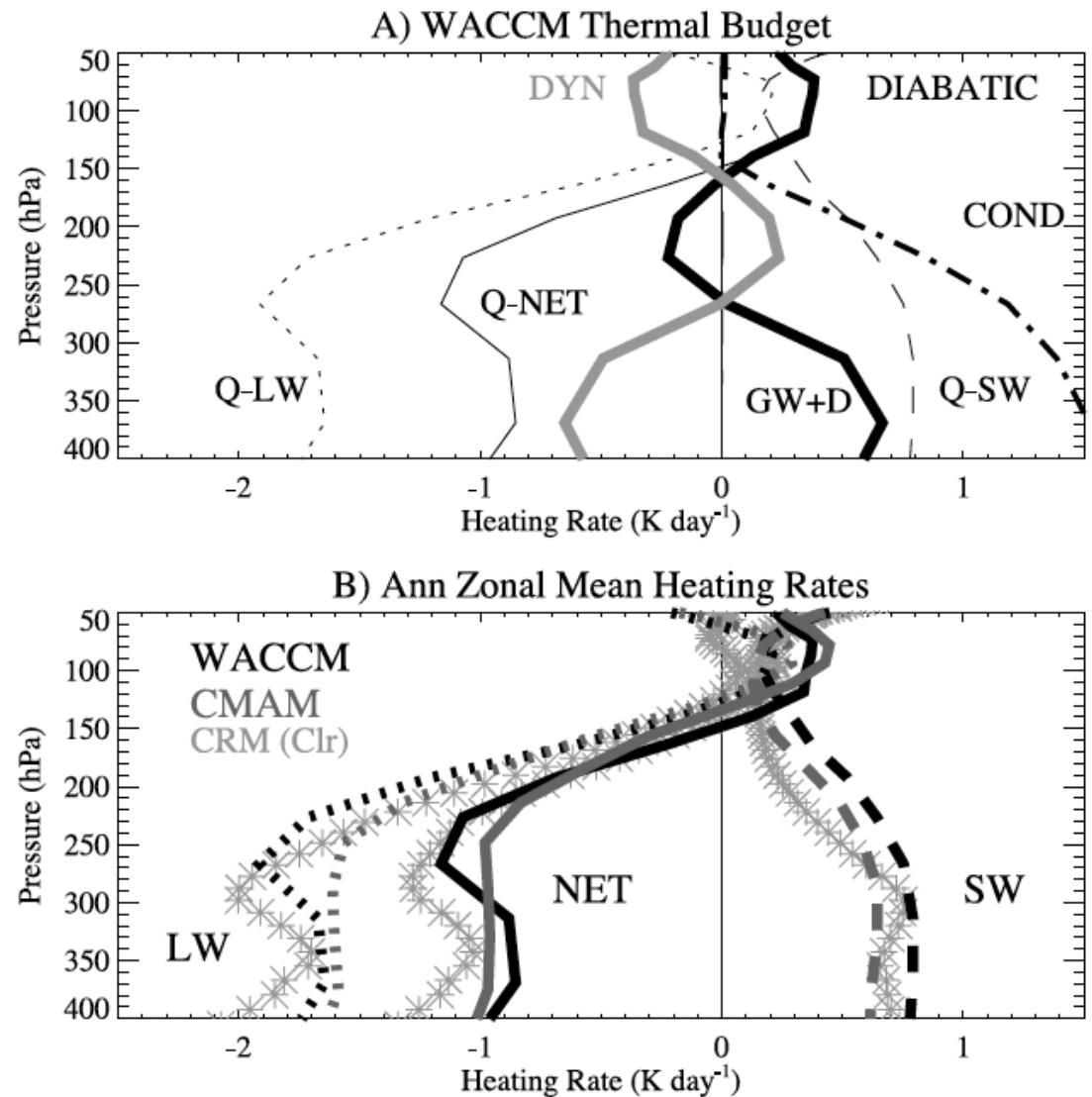


H<sub>2</sub>O has the biggest LW uncertainty: complex radiative transfer

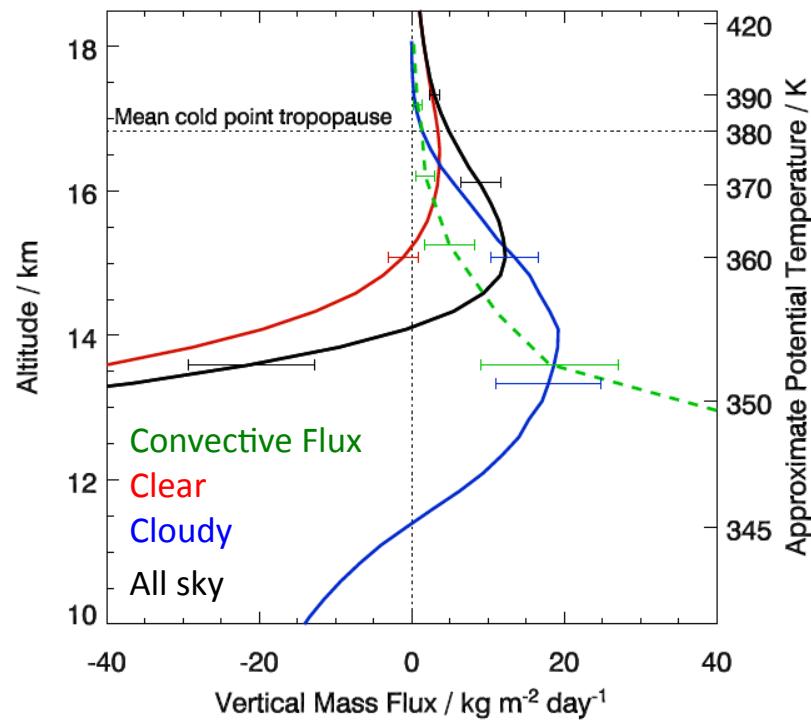
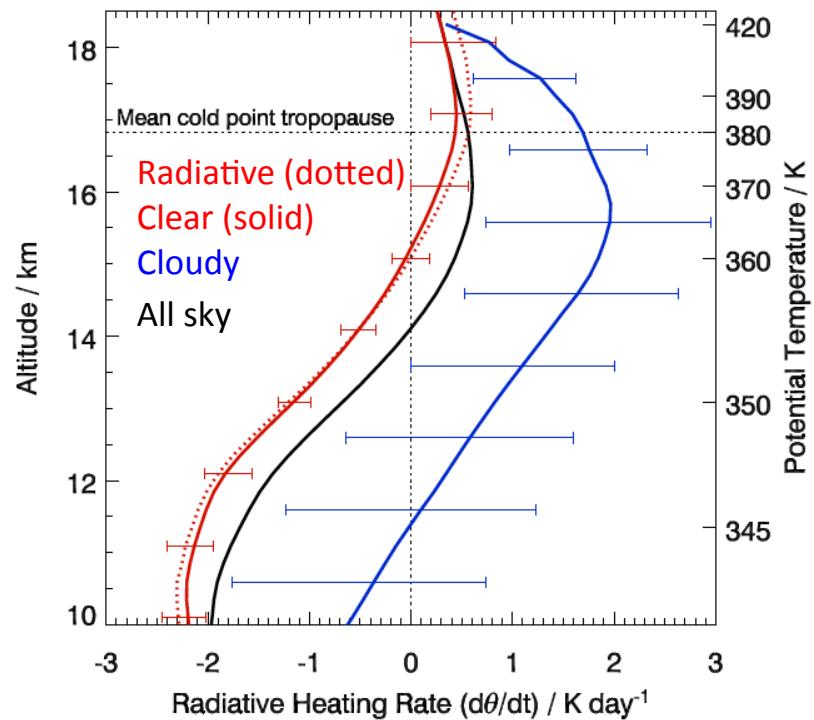
# Role of Dynamics

- Dynamic forcing balanced by diabatic heating/cooling
- Condensation balances LW Radiative Cooling
- Radiation in Models Stable
- Except for water vapor

Gettelman & Birner 2007



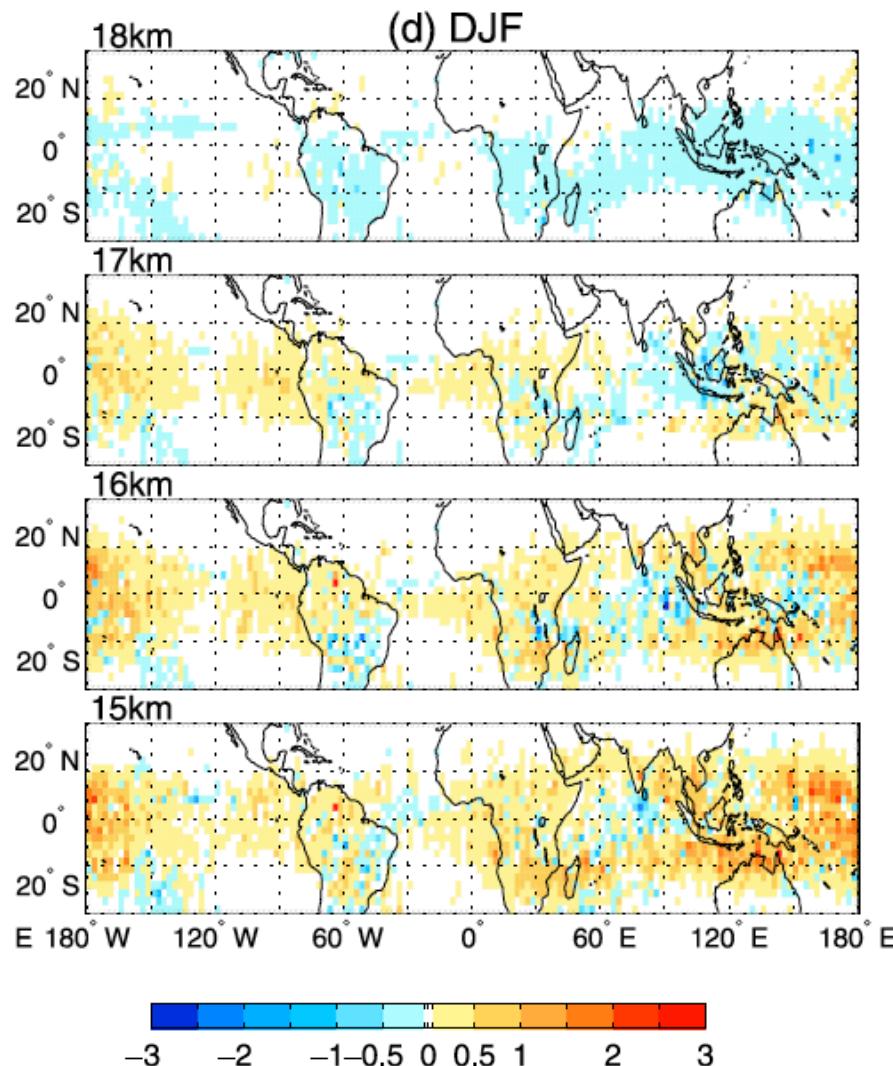
# Impact of Clouds (1)



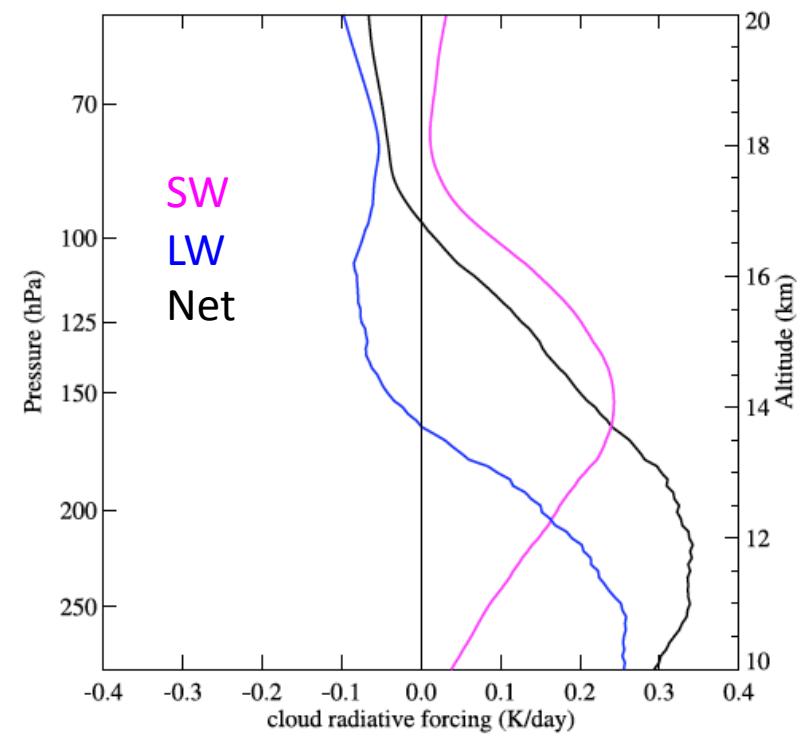
- Cirrus clouds heat TTL (lower  $Q_{clr}=0$ )
- Corti et al 2006, ACP
  - Convective flux from Gettelman et al, 2002

# Clouds (2): Net Heating Rates...

CALIPSO + rad transfer Code



Yang et al 2010, JGR



Thin cirrus ( $t < 0.3$ ) heat at 15-17 km  
~0.5K/d, little effect at 18km  
All sky number similar to Corti et al

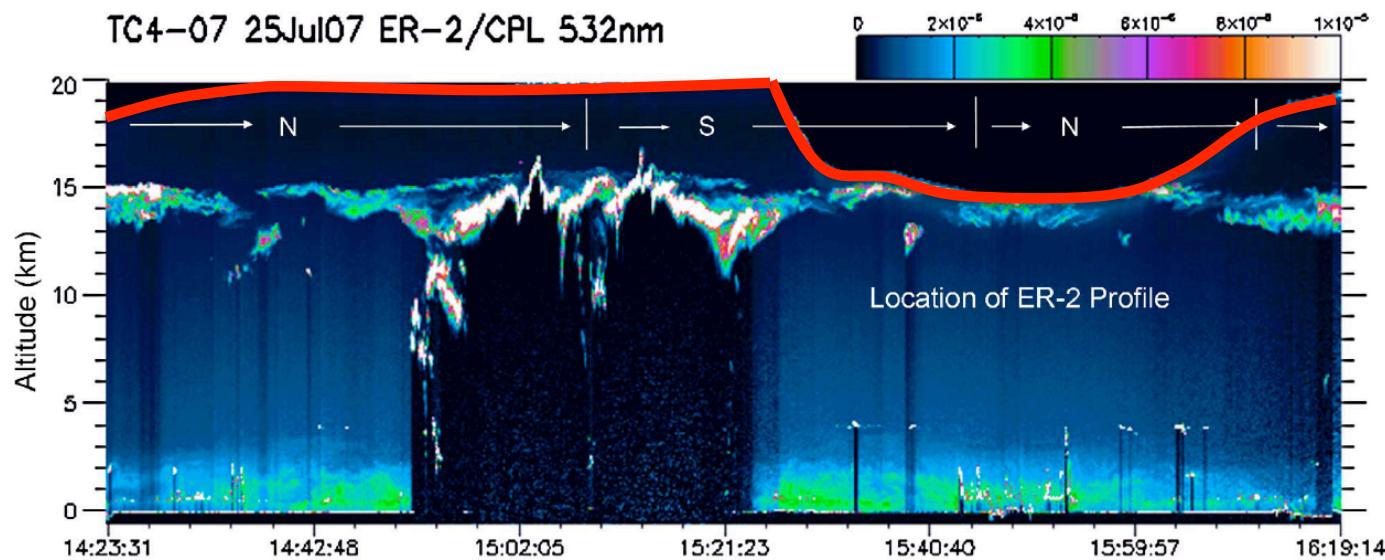
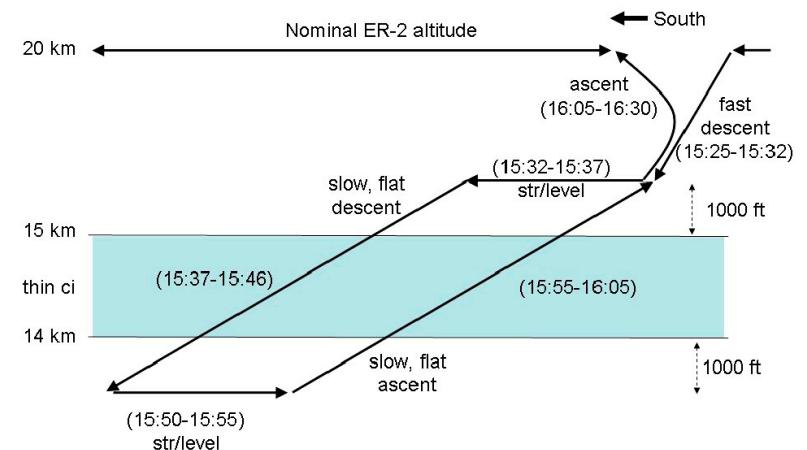
# Clouds (3): TC4 In-Situ heating

Observe divergence of radiative fluxes

Above and below a cloud during TC4

Bucholtz et al 2010

$$\left( \frac{\partial T}{\partial t} \right) = \frac{g}{c_p} \frac{\nabla F}{\Delta p} = 2.5-3.2 \text{ K/d}$$
$$\tau = 0.03$$

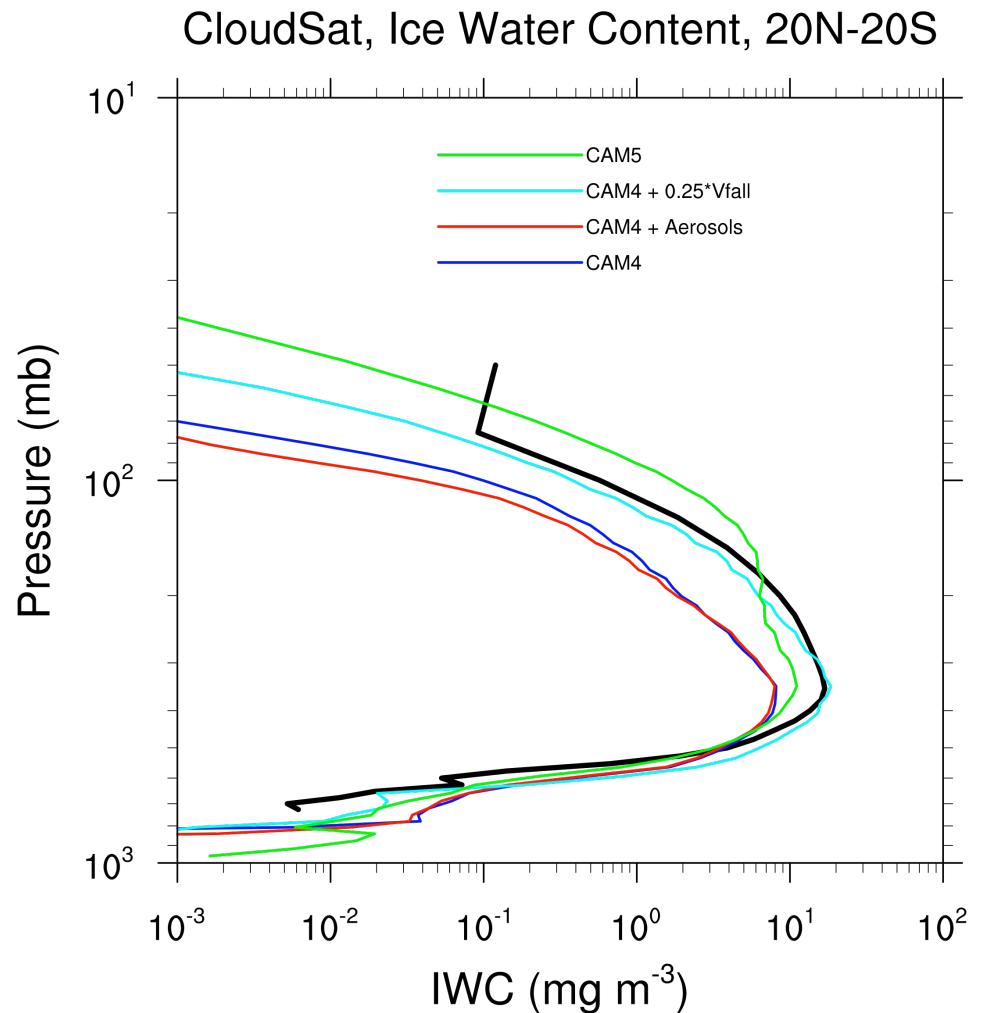


# Model Balances/Quality

- Radiation in clear sky [good]
- Cloud radiative transfer [okay]
- Large scale dynamics [okay]
- Small scale dynamics [bad]
- Cloud climatologies [bad]
- Cloud (cirrus) ice microphysics [ugly]

# Recent CAM/CARMA Results

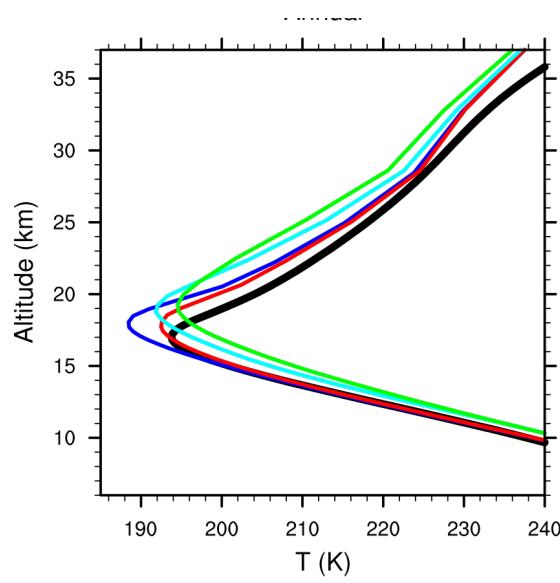
- Black: CloudSat IWC
- Colors: simulations
- Factors of 2 different



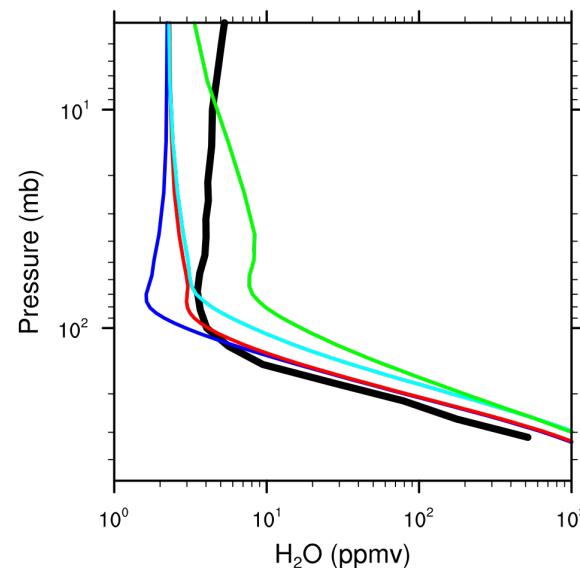
# TTL Temperature & Water Vapor

Cloud differences → different heating, Temps and  $H_2O$

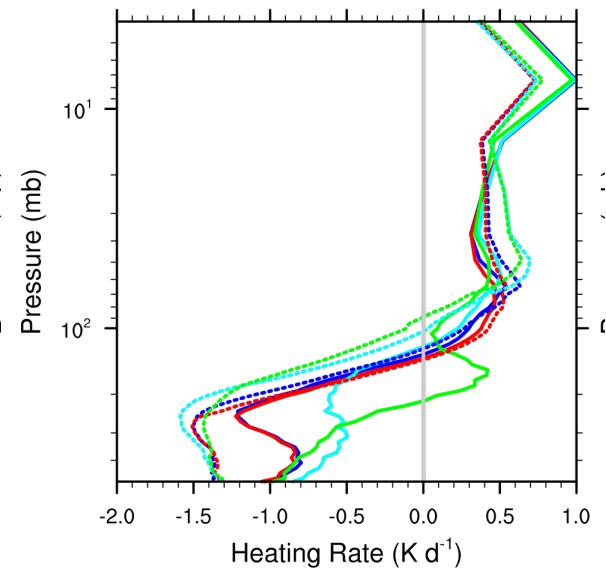
COSMIC T



MLS  $H_2O$



Radiative Heating Rates



CAM5

CAM4 + 0.25\*Vfall

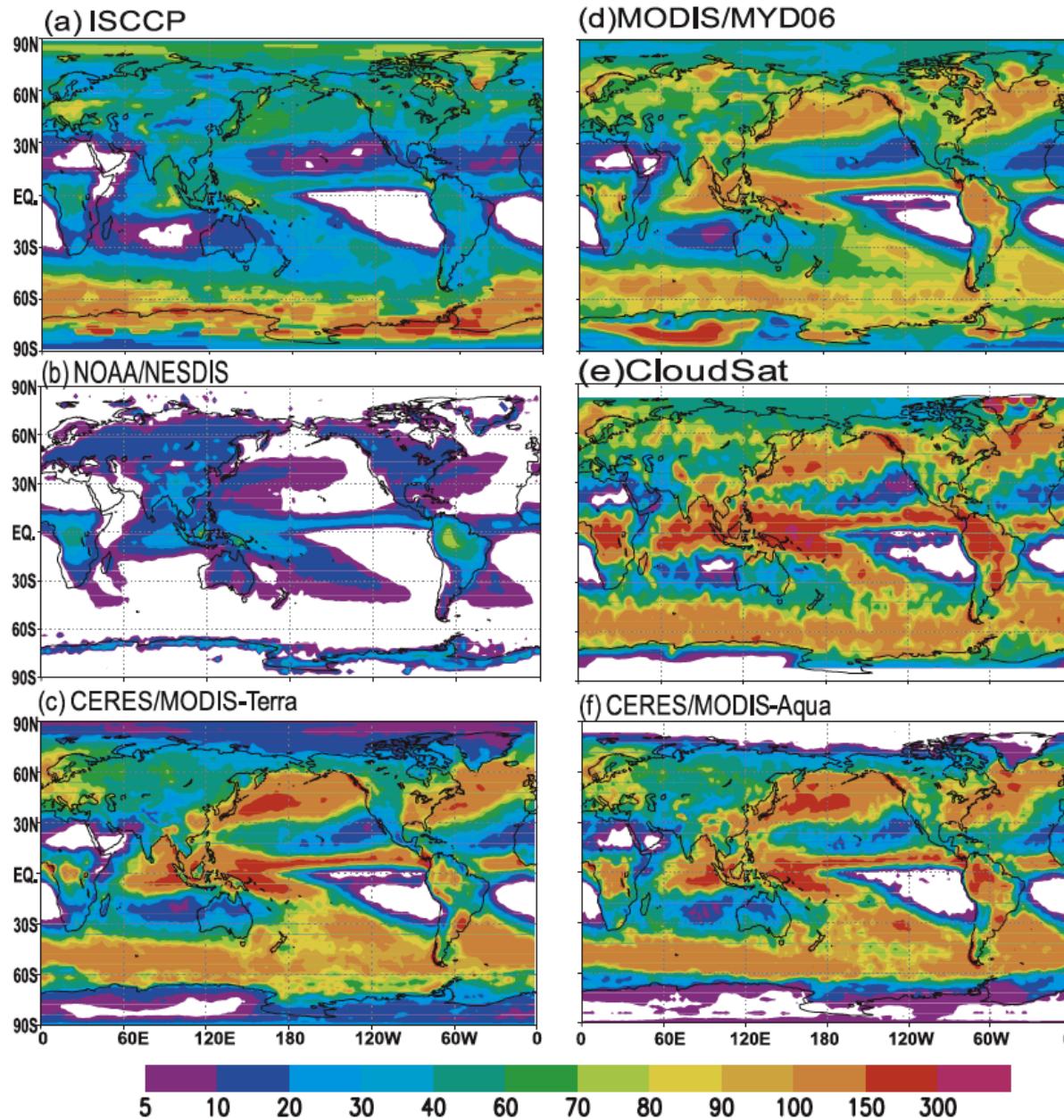
CAM4 + Enhanced Aerosols

CAM4

All Sky (solid)

Clear Sky (dashed)

# IWP Observations



What do we do?  
Adjust clouds, usually  
cloud ice: because  
observations are  
uncertain...

IWP ‘observations’  
from Waliser et al  
2009

Note: large  
differences between  
‘observations’

# What is needed

- More Observed Radiative Fluxes
- Thin Cirrus IWC
  - To evaluate satellites & data sets
  - Biggest uncertainty
  - Frequency on long flight legs
- TTL Temperatures and variability
  - Waves, spectrum
  - Evaluate GPS obs, and below those scales

# Deriving TTL Heating rates

- Vertical structure of radiative fluxes gives ‘clearest’ picture (Bucholtz et al)
- Satellites/Aircraft can do similar things with radiative transfer models
  - Vertical profiles of cloud properties and trace gases help
  - Balloons for vertical range of  $\text{H}_2\text{O}$ ,  $\text{O}_3$ , maybe also a backscatter sonde.
  - Useful at Guam in DJF: have thin cirrus overhead

# ATTREX Flights/Balloons

- Legs through cirrus cloud: heating rates.
  - Aim below 45kft (=13.7 km, aim at 12km)
  - Then turn and ascend to max alt (slow).
- Downward radiometers at max alt
  - Look down at surface and impact of clouds
  - Synergy with NCAR G5 looking up?
- Radiation & clouds on Lagrangian flights
  - Constrain dynamical and radiative heating
- Balloons
  - Profiles of H<sub>2</sub>O, O<sub>3</sub>, Cloud backscatter